UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS

GCE Advanced Subsidiary Level and GCE Advanced Level

MARK SCHEME for the October/November 2010 question paper for the guidance of teachers

9702 PHYSICS

9702/41

Paper 4 (A2 Structured Questions), maximum raw mark 100

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[3]

Page 2	Mark Scheme: Teachers' version	Syllabus	Paper
	GCE AS/A LEVEL – October/November 2010	9702	41

Section A

- 1 (a) force per unit mass (ratio idea essential) B1 [1]
 - (b) graph: correct curvature M1 from $(R, 1.0 g_s)$ & at least one other correct point A1 [2]
 - (c) (i) fields of Earth and Moon are in opposite directions

 either resultant field found by subtraction of the field strength

 or any other sensible comment

 so there is a point where it is zero

 (allow $F_E = -F_M$ for 2 marks)

 M1

 A1

 Equation 1.
 - (ii) $GM_E/x^2 = GM_M/(D-x)^2$ C1 $(6.0 \times 10^{24})/(7.4 \times 10^{22}) = x^2/(60R_E-x)^2$ C1 $x = 54R_E$ A1 [3]
 - (iii) graph: g = 0 at least $\frac{2}{3}$ distance to Moon B1 $g_{\rm E}$ and $g_{\rm M}$ in opposite directions M1 correct curvature (by eye) and $g_{\rm E} > g_{\rm M}$ at surface A1 [3]
- 2 (a) (i) no forces (of attraction or repulsion) between atoms / molecules / particles B1 [1]
 - (ii) sum of kinetic and potential energy of atoms / molecules M1 due to random motion A1 [2]
 - (iii) (random) kinetic energy increases with temperature no potential energy (so increase in temperature increases internal energy)

 A1 [2]
 - (b) (i) zero A1 [1]
 - (ii) work done = $p\Delta V$ C1 = $4.0 \times 10^5 \times 6 \times 10^{-4}$ = 240 J (*ignore any sign*) A1 [2]

(iii)

change	work done / J	heating / J	increase in internal energy / J
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	+240 0 -840	-600 +720 +480	-360 +720 -360

(correct signs essential)
(each horizontal line correct, 1 mark – max 3)

B3

	Page 3		3	Mark Scheme: Teachers' version Syllabus GCE AS/A LEVEL – October/November 2010 9702					
				41					
3	(a)	(i)	reso	nance	B1	[1]			
		(ii)	amp	amplitude 16 mm <u>and</u> frequency 4.6 Hz					
	(b)	(i)	a =	$(-)\omega^2 x$ and $\omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $13.4 \mathrm{m s}^{-2}$	C1 C1 A1	[3]			
		(ii)	F =	<i>ma</i> 150 × 10 ⁻³ × 13.4	C1				
			=	2.0 N	A1	[2]			
	(c)			ys 'below' given line and never zero at 4.6 Hz (or slightly less) and flatter	M1 A1	[2]			
4	(a)	cha	arge /	potential (difference) (ratio must be clear)	B1	[1]			
	(b)	(i)	V =	$Q / 4\pi \varepsilon_0 r$	B1	[1]			
		(ii)	C = so C	$Q/V = 4\pi \varepsilon_0 r$ and $4\pi \varepsilon_0$ is constant $C \propto r$	M1 A0	[1]			
	(c)	(i)	r = ($C / 4\pi \varepsilon_0 r$ $6.8 \times 10^{-12}) / (4\pi \times 8.85 \times 10^{-12})$ 1×10^{-2} m	C1 C1 A1	[3]			
		(ii)	$Q = CV = 6.8 \times 10^{-12} \times 220$ = 1.5 \times 10^{-9} C		A1	[1]			
	(d)	(i)	V = = 83	$Q/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$	A1	[1]			
		(ii)	eithe	er energy = $\frac{1}{2}CV^2$ $\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$	C1 C1				
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = 1.03×10^{-7} J energy = $\frac{1}{2}QV$ $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = 1.03×10^{-7} J	A1 (C1) (C1) (A1)	[3]			

GCE AS/A LEVEL – October/November 2010 9702	41

- 5 (a) field into (the plane of) the paper B1 [1]
 - (b) force due to magnetic field <u>provides</u> the centripetal force $mv^2 / r = Bqv$ C1 $B = (20 \times 1.66 \times 10^{-27} \times 1.40 \times 10^5) / (1.6 \times 10^{-19} \times 6.4 \times 10^{-2})$ B1 = 0.454 T A0 [3]
 - (c) (i) semicircle with diameter greater than 12.8 cm B1 [1]
 - (ii) new flux density = $\frac{22}{20} \times 0.454$ C1 $B = 0.499 \, \text{T}$ A1 [2]
- 6 (a) (i) e.g. prevent flux losses / improve flux linkage B1 [1]
 - (ii) flux in core is changing
 e.m.f. / current (induced) in core
 induced current in core causes heating

 B1
 B1
 - (b) (i) that value of the direct current producing same (mean) power / heating in a resistor M1 [2]
 - (ii) power in primary = power in secondary $V_P I_P = V_S I_S$ A1 [2]
- 7 (a) (i) e.g. electron / particle diffraction B1 [1]
 - (ii) e.g. photoelectric effect B1 [1]
 - **(b) (i)** 6 A1 [1]
 - (ii) change in energy = 4.57×10^{-19} J $\lambda = hc / E$ C1 = $(6.63 \times 10^{-34} \times 3.0 \times 10^{8}) / (4.57 \times 10^{-19})$ = 4.4×10^{-7} m A1 [2]
- 8 (a) splitting of a heavy nucleus (not atom/nuclide) M1 into two (lighter) nuclei of approximately same mass A1 [2]

 - (c) emitted particles have kinetic energy range of particles in the control rods is short / particles stopped in rods / lose kinetic energy in rods B1 kinetic energy of particles converted to thermal energy B1 [3]

Page 5	Mark Scheme: Teachers' version	Syllabus	Paper
	GCE AS/A LEVEL – October/November 2010	9702	41

Section B

9 (a) (i) non-inverting (amplifier)

(ii)
$$(G =) 1 + R_2 / R_1$$

(b) (i) $gain = 1 + 100 / 820$
output = 17 mV

(ii) $9V$
 $(R_2 / R_1 \text{ scores 0 in (a)(ii) but possible 1 mark in each of (b)(i) and (b)(ii)} {1 + R_1 / R_2 \text{ scores 0 in (a)(ii), no mark in (b)(i), possible 1 mark in (b)(ii)} {1 - R_2 / R_1 \text{ or } R_1 / R_2 \text{ scores 0 in (a)(ii), no mark in (b)(i), possible 1 mark in (b)(ii)} {1 - R_2 / R_1 \text{ or } R_1 / R_2 \text{ scores 0 in (a)(ii), ho (ii) and (b)(iii)}}$

10 (a) (i) density × speed of wave (in the medium)

(ii) $\rho = (7.0 \times 10^6) / 4100$
 $= 1700 \text{ kg m}^3$

A1 [1]

(b) (i) $I = I_1 + I_R$

(ii) $1 \cdot \alpha = (0.1 \times 10^6)^2 / (3.1 \times 10^6)^2$
 $= 0.001$

A1 [2]

2. $\alpha \approx 1$

A1 [1]

(c) either very little transmission at an air-skin boundary (almost) complete transmission at a gel-skin boundary when wave travels in or out of the body

or no gel, majority reflection (M1) (M1) when wave travels in or out of the body

11 (a) (i) unwanted random power / signal / energy

(ii) loss of (signal) power / energy

(iii) loss of (signal) power / energy

(b) (i) either signal-to-noise ratio at mic. = $10 \log (P_2 / P_1)$ C1 10

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Pa	ige 6		ark Scheme: Teachers' version /A LEVEL – October/November 2010	Syllabus 9702	Paper 41	•
	(ii) use an am coupled to	nplifier the m		3102	M1 A1	[2]
12 (a)	satellite receiv signal amplifie at a different (different freque e.g. of frequen	es gre d and carrier encies cies u	nitted from Earth to satellite atly attenuated signal transmitted <u>back to Earth</u>) frequency prevent swamping of uplink signal sed (6/4 GHz, 14/11 GHz, 30/20 GHz) any two other for additional physics)	(1) (1) (1) (1)	B1 B1	[4]
(b)	advantage: disadvantage:	e.g. e.g.	much shorter time delay because orbits are much lower whole Earth may be covered in several orbits / with network either must be tracked		M1 A1 (M1) (A1)	
	-	_	or limited use in any one orbit more satellites required for continuous or	peration	M1 A1	[4]