

Markscheme

November 2018

Physics

Higher level

Paper 3



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Section A

| Q | Question | | Answers | Notes | Total |
|----|----------|--|---|--|-------|
| 1. | а | | $m^{\frac{3}{2}} \checkmark$ | Accept other power of tens multiples of $m^{\frac{3}{2}}$, eg: $cm^{\frac{3}{2}}$. | 1 |
| 1. | b | | measured uncertainties «for one oscillation and for 20 oscillations» are the same/similar/OWTTE | | |
| | | | OR | | 2 |
| | | | % uncertainty is less for 20 oscillations than for one ✓ | | |
| | | | dividing «by 20» / finding mean reduces the random error ✓ | | |

(Question 1 continued)

| C | Question | | Answers | Notes | Total |
|----|----------|----|--|---|-------|
| 1. | C | i | Straight line touching at least 3 points drawn across the range \checkmark $ \begin{array}{cccccccccccccccccccccccccccccccccc$ | It is not required to extend the line to pass through the origin. | 1 |
| 1. | С | ii | theory predicts proportional relation $\mbox{$w$} T \propto \frac{1}{d}$, slope = $Td = \frac{c}{\sqrt{g}}$ = constant $\mbox{$w$} \checkmark$ the graph is $\mbox{$w$}$ straight line through the origin $\mbox{$\checkmark$}$ | | 2 |

(Question 1 continued)

| C | uestio | Answers | Notes | Total |
|----|--------|---|---------------------------|-------|
| 1. | d | correctly determines gradient using points where ΔT≥1.5s OR correctly selects a single data point with T≥1.5s ✓ manipulation with formula, any new and correct expression to enable g to be determined ✓ | Allow range 0.51 to 0.57. | Total |
| | | Calculation of g ✓ With g in range 8.6 and 10.7 «m s ⁻² » ✓ | | 4 |

| Q | uestic | on | Answers | Notes | Total |
|----|--------|----|---|-------|-------|
| 2. | а | | to provide a constant heating rate / power | | |
| | | | OR | | 1 |
| | | | to have m proportional to $t \checkmark$ | | |
| 2. | b | | due to heat losses <i>«VIt</i> is larger than heat into liquid» ✓ | | 2 |
| | | | L _v calculated will be larger √ | | 2 |
| 2. | С | | heat losses will be similar / the same for both experiments | | |
| | | | OR | | |
| | | | heat loss presents systematic error ✓ | | |
| | | | taking the difference cancels/eliminates the effect of these losses | | 2 |
| | | | OR | | |
| | | | use a graph to eliminate the effect 🗸 | | |

Section B

Option A — Relativity

| Question | | uestion Answers | | Notes | Total |
|----------|---|-----------------|---|---|-------|
| 3. | а | | a set of rulers and clocks / set of coordinates to record the position and time of events ✓ | | 1 |
| 3. | b | i | the time in frame S' is $t' = \frac{L}{c} \checkmark$ but time is absolute in Galilean relativity so is the same in S \checkmark ALTERNATIVE 2: In frame S, light rays travel at $c + v \checkmark$ so $t = \frac{L}{(c+v)-v} = \frac{L}{c} \checkmark$ | In Alternative 1, they must refer to S' | 2 |
| 3. | b | ii | $x = x' + vt$ and $x' = L$ \checkmark «substitution to get answer» | | 1 |

| C | uestic | on | Answers | Notes | Total |
|----|--------|----|---|-------|-------|
| 4. | а | | $\frac{0.82c + 0.40c}{1 + \frac{0.82c \times 0.40c}{c^2}} \checkmark$ 0.92c \checkmark | | 2 |
| 4. | b | i | $\Delta t' = \frac{120}{0.40c} \checkmark$ $\Delta t' = 1.0 \times 10^{-6} \text{ (s)} \checkmark$ | | 2 |
| 4. | b | | $\gamma = \frac{1}{\sqrt{1 - 0.82^2}} = 1.747 \checkmark$ $\Delta t = \frac{v\Delta x'}{c^2} = 1.747 \times \left(1.0 \times 10^{-6} + \frac{0.82c \times 120}{c^2}\right)$ OR $\Delta t = \frac{120}{1.747 \times (0.92 - 0.82)c} \checkmark$ $2.3 \times 10^{-6} \text{ s.s.} \checkmark$ | | 3 |

| Q | uesti | on | Answers | Notes | Total |
|----|-------|----|--|-------|-------|
| 5. | а | i | $\gamma = \frac{1}{\sqrt{1 - 0.745^2}} = 1.499 \checkmark$ $x' = \frac{1}{\sqrt{1 - 0.745^2}} = 1.499 \times (1.0 - 0) \checkmark$ $x' = 1.5 \text{ m}$ | | 2 |
| 5. | а | ii | $t' = \ll \gamma \left(t - \frac{vx}{c^2} \right) = \gg 1.499 \times \left(0 - \frac{0.745c \times 1}{c^2} \right) \ll - \frac{1.11}{c} \gg$ $\ll ct' = -1.1 \text{ m} \gg$ OR using spacetime interval $0 - 1^2 = (ct')^2 - 1.5^2 \Rightarrow \ll ct' = -1.11 \gg \checkmark$ | | 1 |

(Question 5 continued)

| C | Questi | on | Answers | Notes | Total |
|----|--------|----|--|-------------------------------|-------|
| 5. | b | i | line through event E parallel to <i>ct'</i> axis meeting <i>x'</i> axis and labelled P ✓ | ct S' frame S frame | 1 |
| 5. | b | ii | point on x' axis about $\frac{2}{3}$ of the way to P labelled Q \checkmark | ct S' frame S' frame S frame | 1 |

(Question 5 continued)

| Q | uesti | on | Answers | Notes | Total |
|----|-------|----|--|-----------------------------|-------|
| 5. | С | i | ends of rod must be recorded at the same time in frame S' ✓ any vertical line from E crossing x', no label required ✓ right-hand end of rod intersects at R «whose co-ordinate is less than 1.0 m» ✓ | Ct S' frame 1.0 m S frame | 3 |
| 5. | С | ii | 0.7 m √ | | 1 |

| Q | Question | | Answers | Notes | Total |
|----|----------|----|---|-------|-------|
| 6. | а | | $pc = \sqrt{E^2 - (mc^2)^2} = \sqrt{1.50^2 - 0.511^2} $ «= 1.410 MeV » \checkmark | | 1 |
| 6. | b | i | first equation is due to momentum conservation ✓ second equation is due to total energy conservation ✓ | | 2 |
| 6. | b | ii | adding $2p_1 = 3.42 \text{MeV c}^{-1} \Rightarrow p_1 = 1.71 \text{MeV c}^{-1} \checkmark$ $p_2 = 0.30 \text{MeV c}^{-1} \checkmark$ | | 2 |

| C | Question | | Answers | Notes | Total |
|----|----------|----|---|-------|-------|
| 7. | а | i | the distance from the black hole at which the escape speed is the speed of light ✓ | | 1 |
| 7. | а | ii | $R_{\rm S} = \frac{2GM}{c^2} = \frac{2 \times 6.67 \times 10^{-11} \times 4.0 \times 10^{36}}{9.0 \times 10^{16}} = 3.9 \times 10^9 \text{ m/s} \checkmark$ | | 1 |
| 7. | b | | $2 = \frac{1}{\sqrt{1 - \frac{5.9 \times 10^9}{r}}} \checkmark$ rearranged to give r OR $r = 1.33 \times 5.9 \times 10^9 \text{ mm} \checkmark$ $r = 7.9 \times 10^9 \text{ mm} \checkmark$ | | 3 |

Option B — Engineering physics

| Q | uestic | on | Answers | Notes | Total |
|----|--------|----|---|---|-------|
| 8. | а | | taking torques about the pivot $R \times 4.00 = 36.0 \times 2.5$ \checkmark $R = 22.5$ «N» \checkmark | | 2 |
| 8. | b | i | $36.0 \times 2.50 = 30.6 \times \alpha$ \checkmark $\alpha = 2.94 \text{ «rad s}^{-2} \text{ » } \checkmark$ | | 2 |
| 8. | b | ii | the equation can be applied only when the angular acceleration is constant ✓ any reasonable argument that explains torque is not constant, giving non constant acceleration ✓ | eg weight is no longer perpendicular to the rod | 2 |
| 8. | С | i | «from conservation of energy» Change in GPE = Change in rotational KE \checkmark $W\frac{L}{2} = \frac{1}{2}I\omega^2 \checkmark$ $\omega = \sqrt{\frac{36.0 \times 5.00}{30.6}} \checkmark$ $\omega = 2.4254 \text{ rad s}^{-1}$ » | | 3 |
| 8. | С | ii | $L = 30.6 \times 2.43 = 74.4 \text{ «Js.»}$ | | 1 |

| C | Question | | Answers | Notes | Total |
|----|----------|----|--|-------|-------|
| 9. | а | i | ALTERNATIVE 1: $P_{c} = P_{B} = \frac{P_{A}V_{A}}{V_{B}} \checkmark$ $= \frac{2.8 \times 10^{6} \times 1 \times 10^{-4}}{2.8 \times 10^{-4}} = 1.00 \times 10^{6} Pa $ $ALTERNATIVE 2$ $2.80 \times 10^{6} \times 1.00^{\frac{5}{3}} = P_{c} \times 1.85^{\frac{5}{3}} \checkmark$ $P_{c} = 2.80 \times 10^{6} \times \frac{1.00^{\frac{5}{3}}}{1.85^{\frac{5}{3}}} = 1.00 \times 10^{6} Pa $ | | 2 |
| 9. | a | ii | ALTERNATIVE 1: Since $T_B = T_A$ then $T_C = \frac{V_C T_B}{V_B}$ \checkmark $= \frac{1.85 \times 385}{2.8} \ll = 254.4 \text{K} \gg \checkmark$ ALTERNATIVE 2: $\frac{2.80 \times 1.00}{385} = \frac{1.00 \times 1.85}{T_c} \ll \text{K} \gg \checkmark$ $T_c = 385 \times \frac{1.00 \times 1.85}{2.80} \ll = 254.4 \text{ K} \gg \checkmark$ | | 2 |

(Question 9 continued)

| Q | Question | | Answers | Notes | Total |
|----|----------|----|---|------------------------|-------|
| 9. | b | | work done = $\sqrt{p}\Delta V = 1.00 \times 10^6 \times (1.85 \times 10^{-4} - 2.80 \times 10^{-4}) = \text{»} - 95 \text{ «J» } \checkmark$ change in internal energy = $\sqrt{3} \frac{p}{2} \Delta V = -\frac{3}{2} \times 95 = \text{»} - 142.5 \text{ «J» } \checkmark$ | Allow positive values. | 3 |
| | | | $Q = -95 - 142.5$ \checkmark (-238 J) | | |
| 9. | С | i | net work is $288 - 238 = 50$ «J» ✓ efficiency = « $\frac{288 - 238}{288}$ = » 0.17 ✓ | | 2 |
| 9. | С | ii | along B→C ✓ | | 1 |

| Q | Question | | Answers | Notes | Total |
|-----|----------|----|--|---|-------|
| 10. | а | | ice displaces its own weight of water / OWTTE | | |
| | | | or melted ice volume equals original volume displaced / OWTTE ✓ | | 2 |
| 10. | b | | no change will take place \checkmark continuity equation says $v \times A_1 = u \times A_2$ \checkmark | | |
| 10. | | ' | wand» $A_1 = 4A_2$ \checkmark wgiving result» | | 2 |
| 10. | b | ii | Bernoulli: $\frac{1}{2}\rho v^2 + \rho gH + P_{\text{atm}} = \frac{1}{2}\rho u^2 + 0 + P_{\text{atm}} \text{ w gives } \frac{1}{2} \times 1000 \times \frac{u^2}{16} + 1000 \times 9.8 \times 5.0 = \frac{1}{2} \times 1000 \times u^2 \checkmark$ $u = 10.2 \text{ cm s}^{-1} \text{ w } \checkmark$ | Accept solving directly via conservation of energy. | 2 |

| Q | Question | | Answers | Notes | Total |
|-----|----------|--|--|------------------------------|-------|
| 11. | а | | because the mass and the driver are out of phase «by π» ✓ so upwards ✓ | Justification needed for MP2 | 2 |
| 11. | b | | ALTERNATIVE 1: | | |
| | | | $\ll Q = 2\pi \frac{A_0^2}{A_0^2 - A_1^2} \Rightarrow \frac{A_1^2}{A_0^2} = 1 - \frac{2\pi}{Q} \checkmark$ | | |
| | | | $\frac{A_1}{A_0} = \sqrt{1 - \frac{2\pi}{22}} = A_1 = 8.5 \text{ cm}$ | | 2 |
| | | | ALTERNATIVE 2: | | |
| | | | driver amplitude is constant ✓ | | |
| | | | so mass amplitude is unchanged at 10 cm ✓ | | |

Option C — Imaging

| Q | Question | | Answers | Notes | Total |
|-----|----------|----|---|---|-------|
| 12. | а | | each incident ray shown splitting into two ✓ each pair symmetrically intersecting each other on principal axis ✓ for red, intersection further to the right ✓ | For MP3, at least one of the rays must be labelled. | 3 |
| 12. | b | i | rays diverge after passing through lens OR the extension of the rays will intersect the principal axis on the side of incident rays/as if they were coming from the focal point/points in the left side/OWTTE ✓ | | 1 |
| 12. | b | ii | by placing a diverging lens next to the converging lens OR make an achromatic doublet ✓ | Further details are not required. | |

| Q | uestic | n Answers | Notes | Total |
|-----|--------|--|--|-------|
| 13. | а | proper construction lines ✓ | eyepiece lens | |
| | | image at intersection of proper construc | tion lines ✓ | |
| | | | objective lens for the state of the state o | 2 |

(Question 13 continued)

| Question | | Answers | Notes | Total |
|----------|----|---|-------|-------|
| 13. b | i | distance of intermediate image from objective is $\frac{1}{v} = \frac{1}{20} - \frac{1}{24} ie: \ v = 120 \text{ mm} \text{ m} \checkmark$ distance of intermediate image from eyepiece is $\frac{1}{u} = \frac{1}{60} - \left(-\frac{1}{240}\right) ie: \ u = 48 \text{ mm} \checkmark$ lens separation 168 mm $ \checkmark$ | | 3 |
| 13. b | ii | ALTERNATIVE 1: eyepiece: $m = \frac{-v}{u} = \frac{240}{48} = 5$ AND objective $m = \frac{-v}{u} = \frac{-120}{24} = -5$ Total $m = -5 \times 5 = -25$ ALTERNATIVE 2: $m = \left(\frac{240}{60} + 1\right) \times \left(-\frac{120}{24}\right)$ $m = -25$ | | 2 |

| Q | Question | | Answers | Notes | Total |
|-----|----------|-----|---|--|-------|
| 14. | а | i | | | 2 |
| 14. | а | ii | to have a critical angle close to 90° ✓ so only rays parallel to the axis are transmitted ✓ to reduce waveguide/modal dispersion ✓ | Do not accept "so that most rays are reflected". | 1 max |
| 14. | b | i | long path is $\frac{12 \times 10^3}{\sin 84^\circ}$ \checkmark = 12066 «m» \checkmark «so 66 m longer» | | 2 |
| 14. | b | ii | speed of light in core is $\frac{3.0 \times 10^8}{1.52} = 1.97 \times 10^8 \text{ wm s}^{-1} \text{ wm s}^{-1}$ where $\frac{66}{1.97 \times 10^8} = 3.35 \times 10^{-7} \text{ ws } \text{ wm s}^{-1}$ | | 2 |
| 14. | b | iii | no, period of signal is 1×10 ⁻⁸ « s » which is smaller than the time delay/OWTTE ✓ | | 1 |

| Question | | ion | Answers | Notes | Total |
|----------|---|-----|--|-------|-------|
| 15. | а | | protons spin direction changes | | |
| | | | OR | | 1 |
| | | | proton energy state changes ✓ | | |
| 15. | b | | Relaxation time «of signal/proton spin» √ | | 2 |
| | | | Location/time delay of the emitted RF signal ✓ | | 2 |
| 15. | С | | Relaxation time gives information on tissue type/density/health/OWTTE✓ | | |
| | | | | | 2 |
| | | | Location information provides 3D image/OWTTE✓ | | |

| Q | uestion | Answers | Notes | Total |
|-----|---------|--|-------|-------|
| 16. | а | $I_0 e^{-23 \times 0.041} \checkmark$ | | |
| | | $= 0.39 I_0 $ / | | 2 |
| 16. | b | $R = \left(\frac{6.3 \times 10^6 - 1.7 \times 10^6}{6.3 \times 10^6 + 1.7 \times 10^6}\right)^2 = 0.33 $ | | 2 |
| | | so reflected intensity is $0.33 \times 0.39 I_0 = 0.13 I_0$ \checkmark | | |
| 16. | С | $0.13I_0 \times 0.39 = 0.05I_0$ \checkmark | | 1 |

Option D — Astrophysics

| Q | Question | | Answers | Notes | Total |
|-----|----------|----|---|-------|-------|
| 17. | а | | In cluster, stars are gravitationally bound <i>OR</i> constellation not ✓ | | |
| | | | In cluster, stars are the same/similar age <i>OR</i> in constellation not ✓ | | |
| | | | Stars in cluster are close in space/the same distance | | |
| | | | OR | | 2 max |
| | | | in constellation not ✓ | | |
| | | | Cluster stars appear closer in night sky than constellation ✓ | | |
| | | | Clusters originate from same gas cloud <i>OR</i> constellation does not ✓ | | |
| 17. | b | i | d=275 «pc» ✓ | | 1 |
| 17. | b | ii | because of the difficulty of measuring very small angles ✓ | | 1 |
| 17. | С | | mass of gas cloud > Jeans mass √ | | |
| | | | «magnitude of» gravitational potential energy > E _k of particles ✓ | | 2 max |
| | | | cloud collapses/coalesces «to form a protostar» ✓ | | |

| Q | uestic | on | Answers | Notes | Total |
|-----|--------|-----|---|--|-------|
| 18. | а | i | $\lambda = \frac{2.9 \times 10^{-3}}{4600} = 800 \text{ mm}$ | | 1 |
| 18. | а | ii | black body curve shape ✓ peaked at a value from range 600 to 660 nm ✓ | | 2 |
| 18. | а | iii | $\frac{L}{L_{\odot}} = \left(\frac{0.73R_{\odot}}{R_{\odot}}\right)^{2} \times \left(\frac{4600}{5800}\right)^{4} \checkmark$ $L = 0.211L_{\odot} \checkmark$ | | 2 |
| 18. | b | | $M = \text{«} 0.21^{\frac{1}{3.5}} M_{\odot} = \text{»} 0.640 M_{\odot} \checkmark$ | Accept reverse argument 0.64 ^{3.5} = 0.21 | 1 |
| 18. | С | | $\frac{T_E}{T_\odot} = \frac{\frac{M_E}{L_E}}{\frac{M_\odot}{L_\odot}} = \frac{0.64}{0.21} = 3.0 \checkmark$ $T \approx 27 \text{ billion years } \checkmark$ | | 2 |
| 18. | d | | red giant ✓ planetary nebula ✓ white dwarf ✓ | do NOT accept supernova, red supergiant, neutron star or black hole as stages | 3 |

| Q | Question | | Answers | Notes | Total |
|-----|----------|--|--|-------|-------|
| 19. | а | | measured redshift «z» of star \checkmark use of Doppler formula OR z~v/c OR $v = \frac{c\Delta\lambda}{\lambda}$ to find $v \checkmark$ | OWTTE | 2 |
| 19. | b | | use of gradient or any point on the line to obtain any expression for either $H = \frac{V}{d}$ or $t = \frac{d}{V}$ correct conversion of d to m and v to m/s \checkmark = 4.6×10^{17} «s» \checkmark | | 3 |

| Q | uestion | Answers | Notes | Total |
|-----|---------|--|-------|-------|
| 20. | а | energy filling all space ✓ resulting in a repulsive force/force opposing gravity ✓ accounts for the accelerating universe ✓ makes up about 70% of «the energy» of universe ✓ | | 2 max |
| 20. | b | black hole ✓ brown dwarf ✓ massive compact halo object /MACHO✓ neutrinos ✓ weakly interacting massive particle /WIMP ✓ | | 2 max |

| Q | Question | | Answers | Notes | Total |
|-----|----------|--|---|-------|-------|
| 21. | а | | «wavelength of light/CBR» $\lambda \propto R$ \checkmark | OWTTE | |
| | | | reference to Wien's law showing that $\lambda \propto \frac{1}{T} \checkmark$ | | 3 |
| | | | combine to get result √ | | |
| 21. | b | | $\frac{R_{\text{past}}}{R_{\text{now}}} = \frac{3}{300} = 0.01 \ \checkmark$ | | 1 |