

**OXFORD CAMBRIDGE AND RSA EXAMINATIONS
GCSE TWENTY FIRST CENTURY SCIENCE**

A183/01

**PHYSICS A/FURTHER ADDITIONAL
SCIENCE A**

Module P7 (Foundation Tier)

MONDAY 23 JUNE 2014: Morning

**DURATION: 1 hour
plus your additional time allowance**

MODIFIED ENLARGED

Candidate forename		Candidate surname	
-------------------------------	--	------------------------------	--

Centre number						Candidate number				
--------------------------	--	--	--	--	--	-----------------------------	--	--	--	--

Candidates answer on the Question Paper.

A calculator may be used for this paper.

OCR SUPPLIED MATERIALS:

None

OTHER MATERIALS REQUIRED:

Pencil

Ruler (cm/mm)

READ INSTRUCTIONS OVERLEAF

INSTRUCTIONS TO CANDIDATES

Write your name, centre number and candidate number in the boxes on the first page. Please write clearly and in capital letters.

Use black ink. HB pencil may be used for graphs and diagrams only.

Answer ALL the questions.

Read each question carefully. Make sure you know what you have to do before starting your answer.

Write your answer to each question in the space provided. Additional paper may be used if necessary but you must clearly show your candidate number, centre number and question number(s).

INFORMATION FOR CANDIDATES

The quality of written communication is assessed in questions marked with a pencil (✎).

A list of useful relationships is printed on pages 3 to 5.

The number of marks is given in brackets [] at the end of each question or part question.

The total number of marks for this paper is 60.

Any blank pages are indicated.

TWENTY FIRST CENTURY SCIENCE EQUATIONS

USEFUL RELATIONSHIPS

THE EARTH IN THE UNIVERSE

$$\text{distance} = \text{wave speed} \times \text{time}$$

$$\text{wave speed} = \text{frequency} \times \text{wavelength}$$

SUSTAINABLE ENERGY

$$\text{energy transferred} = \text{power} \times \text{time}$$

$$\text{power} = \text{voltage} \times \text{current}$$

$$\text{efficiency} = \frac{\text{energy usefully transferred}}{\text{total energy supplied}} \times 100\%$$

EXPLAINING MOTION

$$\text{speed} = \frac{\text{distance travelled}}{\text{time taken}}$$

$$\text{acceleration} = \frac{\text{change in velocity}}{\text{time taken}}$$

$$\text{momentum} = \text{mass} \times \text{velocity}$$

$$\frac{\text{change of momentum}}{\text{}} = \frac{\text{resultant force}}{\text{}} \times \frac{\text{time for which it acts}}{\text{}}$$

$$\frac{\text{work done by a force}}{\text{}} = \text{force} \times \frac{\text{distance moved in the direction of the force}}{\text{}}$$

$$\text{amount of energy transferred} = \text{work done}$$

$$\frac{\text{change in gravitational potential energy}}{\text{}} = \text{weight} \times \frac{\text{vertical height difference}}{\text{}}$$

$$\text{kinetic energy} = \frac{1}{2} \times \text{mass} \times [\text{velocity}]^2$$

ELECTRIC CIRCUITS

$$\text{power} = \text{voltage} \times \text{current}$$

$$\text{resistance} = \frac{\text{voltage}}{\text{current}}$$

$$\frac{\text{voltage across primary coil}}{\text{voltage across secondary coil}} = \frac{\text{number of turns in primary coil}}{\text{number of turns in secondary coil}}$$

RADIOACTIVE MATERIALS

$$\text{energy} = \text{mass} \times [\text{speed of light in a vacuum}]^2$$

OBSERVING THE UNIVERSE

$$\text{lens power} = \frac{1}{\text{focal length}}$$

$$\text{magnification} = \frac{\text{focal length of objective lens}}{\text{focal length of eyepiece lens}}$$

$$\text{speed of recession} = \text{Hubble constant} \times \text{distance}$$

$$\text{pressure} \times \text{volume} = \text{constant}$$

$$\frac{\text{pressure}}{\text{temperature}} = \text{constant}$$

$$\frac{\text{volume}}{\text{temperature}} = \text{constant}$$

$$\text{energy} = \text{mass} \times [\text{speed of light in a vacuum}]^2$$

Answer ALL the questions.

1 Much of our knowledge about stars comes from the light we receive from the stars.

(a) (i) It is very useful to produce a spectrum of the light from a star.

Draw a labelled diagram to show how a prism can produce a spectrum from a beam of light.

[3]

- (ii) What is the name of the process in the prism that causes the spectrum?

Put a **ring** around the correct answer.

absorption

parallax

reflection

refraction

[1]

- (iii) What else can be used to produce a spectrum, other than a prism?

_____ [1]

(b) The spectrum from a star can give important information about the star.

Complete the sentences about lines in a spectrum.

Choose the best words from the list.

electrons

light

lines

neutrons

sound

Dark lines in the spectrum are produced when

_____ is absorbed.

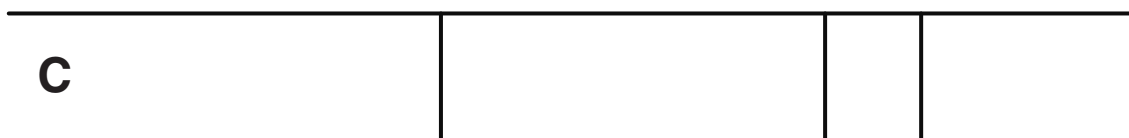
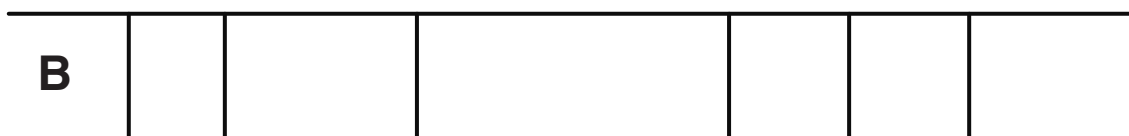
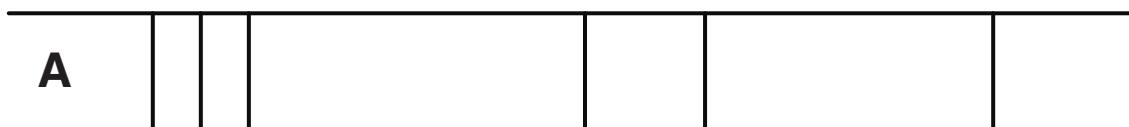
Each element gives a different pattern of

_____ in a star's spectrum.
[2]

(c) (i) The diagram shows the spectrum from a star



Use the spectra for the elements A, B, C, D shown below to work out which TWO elements are in the star.



Which TWO elements A, B, C or D are in the star?

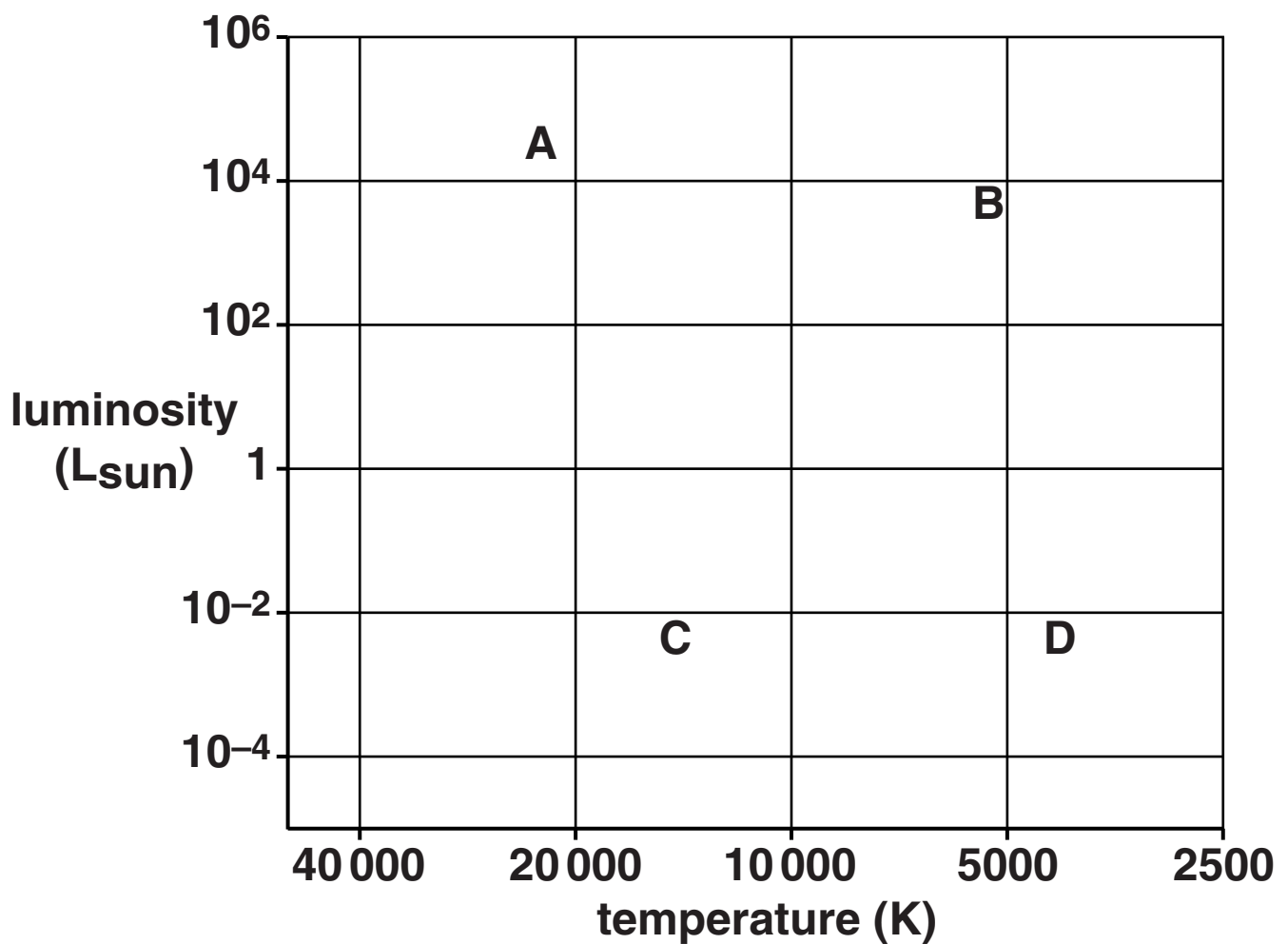
answer _____ and _____ [2]

(ii) What are the NAMES of the two most common elements in a young star?

_____ and _____ [2]

[TOTAL: 11]

2 This question is about the Hertzsprung-Russell diagram.



(a) The Hertzsprung-Russell diagram above shows the temperature and brightness of four objects A, B, C and D.

Which object on the Hertzsprung-Russell diagram will be:

hottest and brightest? A, B, C or D? _____

coolest and dimmest? A, B, C or D? _____

[2]

(b) What type of object is the Hertzsprung-Russell diagram designed to show?

Put a ring around the correct answer.

galaxies

moons

planets

stars

[1]

(c) (i) The Sun has a temperature of about 5000 K and luminosity of $1 L_{\text{sun}}$.

Plot and label the position of the Sun on the Hertzsprung-Russell diagram opposite. [2]

(ii) The average temperature of the Earth is about 15°C .

What is this temperature in kelvin (K)?

temperature = _____ K [2]

**(iii) The Earth can not be plotted on this Hertzsprung-Russell diagram.
Why not?**

_____ [1]

(d) A, C and D on the Hertzsprung-Russell diagram are stars.

Draw straight lines to connect the STARS on the left to the TYPE OF STAR on the right.

STARS

A and D

C

TYPE OF STAR

main sequence

red giant

supergiant

white dwarf

[2]

[TOTAL: 10]

3 Describe the life of a star like the Sun, from its formation to its death.



The quality of written communication will be assessed in your answer.

[6]

[TOTAL: 6]

- 4 The table gives some information about possible sites for a new astronomical observatory.

Site	Height above sea level in m	Average cloudless nights per year	Distance to nearest town in km
W	5000	360	100
X	1000	120	150
Y	6000	270	50
Z	500	230	30

Which site would be the best for an astronomical observatory?

By considering each site, explain and justify your choice.



The quality of written communication will be assessed in your answer.

[6]

[TOTAL: 6]

5 In recent years, scientists have observed things outside our solar system. These observations make them think it is more likely that life will be found outside our solar system.

(a) What have scientists observed and how do these observations make it more likely that life will be found outside the solar system?

[2]

(b) How many species of extraterrestrial life have scientists found?

[1]

[TOTAL: 3]

6 Draw a labelled diagram to explain a LUNAR eclipse.

[3]

[TOTAL: 3]

7 In the late 1700s, the Titius-Bode Law was published.

The law was used for calculating the distance of the planets from the Sun. The distance from the Earth to the Sun is 1 AU. This is what the law says:

To find the distance in AU:

**take the sequence of numbers 0, 3, 6, 12, 24, 48, 96, ...
(each number after the first two is double the previous number)**

add 4 to each number in the sequence

divide each number by 10 to give the distance.

Titius-Bode distance calculation in AU	Planet	Actual distance from Sun in AU
$(0 + 4)/10 = 0.4$	Mercury	0.39
$(3 + 4)/10 = 0.7$	Venus	0.72
$(6 + 4)/10 = 1.0$	Earth	1.00
$(12 + 4)/10 = 1.6$	Mars	1.52
$(48 + 4)/10 = 5.2$	Jupiter	5.20
$(96 + 4)/10 = 10$	Saturn	9.54

(a) Suggest why the Titius-Bode Law was only applied to the six planets out to Saturn in the first instance.

[1]

(b) Bode thought there should be a planet between Mars and Jupiter.

(i) Calculate the distance using the Titius-Bode Law.

distance = _____ AU [2]

(ii) In 1801, the astronomer Giuseppe Piazzi discovered a new (dwarf) planet, *Ceres*, at a distance of 2.77 AU from the Sun.

**Does this support the Titius-Bode Law?
Explain why.**

_____ [1]

(iii) Giuseppe Piazzi did not make enough observations to describe the orbit of *Ceres*. Other astronomers could not find the planet.

Why is it important that other astronomers observe the new planet?

_____ [2]

(c) The table below gives the data for other more recently discovered planets.

Titius-Bode calculation of distance in AU	Planet	Actual distance from Sun in AU
$(192 + 4)/10 = 19.6$	Uranus	19.18
$(384 + 4)/10 = 38.8$	Neptune	30.06
$(768 + 4)/10 = 77.2$	Pluto	39.44

Discuss how these results affect confidence in the Titius-Bode Law.

[4]

[TOTAL: 10]

BLANK PAGE

- 8 One of the most distant objects visible to the naked eye is the Andromeda galaxy.**

Edwin Hubble first measured the distance to the Andromeda galaxy using Cepheid variables. He measured the distance as about 1 million light years. Modern measurements using Cepheid variables, give a distance of 2.5 million light years.

Telescopes in space have made it possible to make better measurements of parallax and of the brightness of stars.

Explain:

how using space telescopes gives better measurements of parallax and brightness

how this improves measurement of distance to Cepheid variables.



The quality of written communication will be assessed in your answer.

[6]

[Total: 6]

- 9 Ian is a young scientist at university. He reads about some research in a peer reviewed scientific journal.**

(a) What is meant by ‘peer reviewed’?

[2]

- (b) The research measured the speed of recession of a galaxy.
Ian’s supervisor asks him to reproduce the results.**

- (i) Ian looks up the distance to the galaxy in a reference book and then calculates the speed of recession for the galaxy using the Hubble constant.**

distance to galaxy = 2000 Mpc

Hubble constant = 70 km/s per Mpc

Calculate the speed of recession of the galaxy.

speed of recession = _____ km/s [2]

(ii) Ian's results agree with the published research.

However Ian's supervisor tells him that he has not reproduced the results.

What is the problem with Ian's method?

[1]

[TOTAL: 5]

END OF QUESTION PAPER

BLANK PAGE

BLANK PAGE



Copyright Information

OCR is committed to seeking permission to reproduce all third-party content that it uses in its assessment materials. OCR has attempted to identify and contact all copyright holders whose work is used in this paper. To avoid the issue of disclosure of answer-related information to candidates, all copyright acknowledgements are reproduced in the OCR Copyright Acknowledgements Booklet. This is produced for each series of examinations and is freely available to download from our public website (www.ocr.org.uk) after the live examination series.

If OCR has unwittingly failed to correctly acknowledge or clear any third-party content in this assessment material, OCR will be happy to correct its mistake at the earliest possible opportunity.

For queries or further information please contact the Copyright Team, First Floor, 9 Hills Road, Cambridge CB2 1GE.

OCR is part of the Cambridge Assessment Group; Cambridge Assessment is the brand name of University of Cambridge Local Examinations Syndicate (UCLES), which is itself a department of the University of Cambridge.

